
Computer Graphics III

Spherical integrals, Light & Radiometry – Exercises

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Reminders & org

- **Renderings due next week**
 - Upload to google drive, show on the big screen, 5 minutes per team (how many teams do we have)
- Papers for presentations in the lab – 7.11., 21.11,
 - ACM TOG special issue on production rendering
<https://dl.acm.org/citation.cfm?id=3243123&picked=prox>
- Reminder – choose papers for the exam
 - <http://kesen.realtimerendering.com/>
- Log your choices here
 - <https://docs.google.com/document/d/128e4DghoIvH64DI6Ohu2eRGthom5i8WlKpDwNyJzpVM/edit?usp=sharing>
- Decide assignments track vs. individual project track by Wed, Oct 31st 2018.



PEN & PAPER EXERCISES

Surface area of a (subset of a) sphere

- Calculate the surface area of a unit sphere.
- Calculate the surface area of a spherical cap delimited by the angle θ_0 measured from the north pole.
- Calculate the surface area of a spherical wedge with angle ϕ_0 .

Solid angle

- What is the solid angle under which we observe an (infinite) plane from a point outside of the plane?
- Calculate the solid angle under which we observe a sphere with radius R , the center of which is at the distance D from the observer.

Isotropic point light

- **Q:** What is the emitted power (flux) of an isotropic point light source with intensity that is a constant I in all directions?

Isotropic point light

- **A:** Total flux:

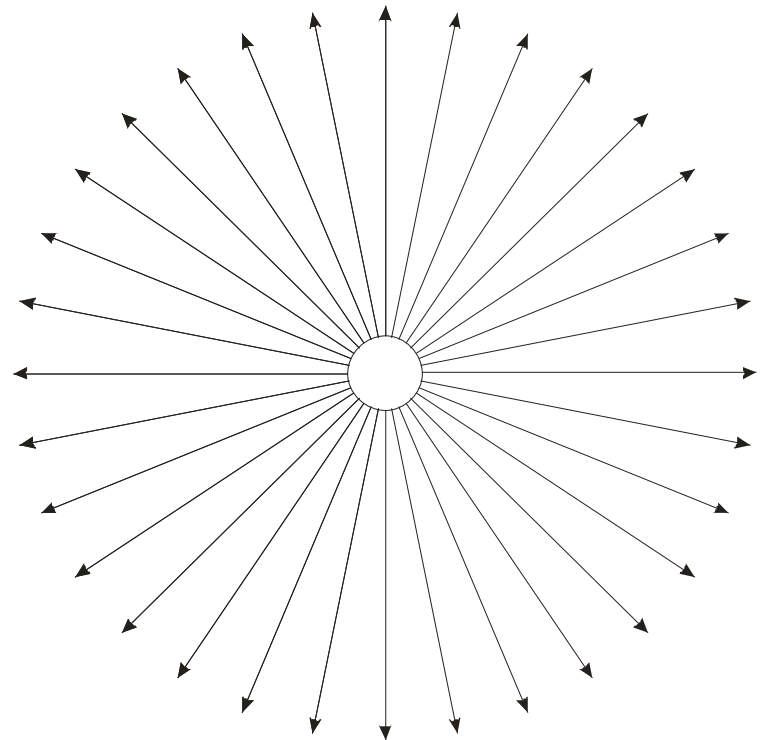
$$\Phi = \int_{\Omega} I(\omega) d\omega = \left| \begin{array}{l} \text{substitute:} \\ d\omega = \sin \theta d\theta d\varphi \end{array} \right|$$

$$= I \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi} \sin \theta d\theta d\varphi$$

$$= I 2\pi \left[-\cos \theta \right]_0^{\pi}$$

$$= 4\pi I$$

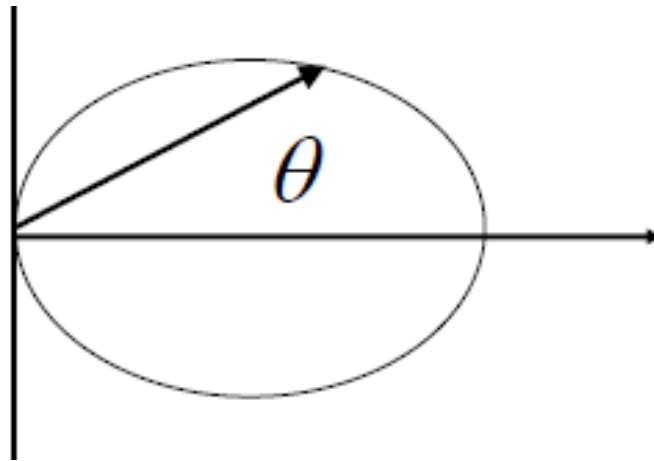
$$I = \frac{\Phi}{4\pi}$$



Cosine spot light

- What is the power (flux) of a point source with radiant intensity given by:

$$I(\omega) = I_0 \max\{0, \omega \cdot \vec{d}\}^s$$



Spotlight with linear angular fall-off

- What is the power (flux) of a point light source with radiant intensity given by:

$$I(\theta, \phi) = \begin{cases} I_0 & \theta \leq \alpha \\ I_0 \frac{\beta - \theta}{\beta - \alpha} & \alpha < \theta < \beta \\ 0 & \theta \geq \beta \end{cases}$$

Constant part

$$\Phi_1 = \int_0^{2\pi} \int_0^\alpha I_0 \sin \theta d\theta d\phi = I_0 2\pi (1 - \cos \alpha).$$

Linear part

$$\Phi_2 = \int_0^{2\pi} \int_\alpha^\beta I_0 \frac{\beta - \theta}{\beta - \alpha} \sin \theta d\theta d\phi = I_0 \frac{2\pi}{\beta - \alpha} \int_\alpha^\beta (\beta - \theta) \sin \theta d\theta \quad (1)$$

The last integral is the sum of the following two integrals:

$$\int_\alpha^\beta \beta \sin \theta d\theta = \beta \cos \alpha - \beta \cos \beta \quad (2)$$

$$- \int_\alpha^\beta \theta \sin \theta d\theta = \left| \sin \theta - \theta \cos \theta \right|_\beta^\alpha = \sin \alpha - \alpha \cos \alpha - \sin \beta + \beta \cos \beta \quad (3)$$

Plugging (2) and (3) into (1) and rearranging, we get

$$\Phi_2 = I_0 \frac{2\pi}{\beta - \alpha} [(\beta - \alpha) \cos \alpha + \sin \alpha - \sin \beta] = I_0 2\pi \left[\cos \alpha - \frac{\sin \beta - \sin \alpha}{\beta - \alpha} \right]. \quad (4)$$

Total flux

$$\Phi = \Phi_1 + \Phi_2 = I_0 2\pi \left[1 - \frac{\sin \beta - \sin \alpha}{\beta - \alpha} \right] \quad (5)$$

Irradiance due to a Lambertian light source

- What is the irradiance $E(\mathbf{x})$ at point \mathbf{x} due to a uniform Lambertian area source observed from point \mathbf{x} under the solid angle Ω ?

How dark are outdoor shadows?

- ◆ luminance arriving on a surface from a full (overhead) sun is $300,000 \times$ luminance arriving from the blue sky, but the sun occupies only a small fraction of the sky
- ◆ illuminance on a sunny day = 80% from the sun + 20% from blue sky, so shadows are $1/5$ as bright as lit areas (2.3 f/stops)

(Marc Levoy)

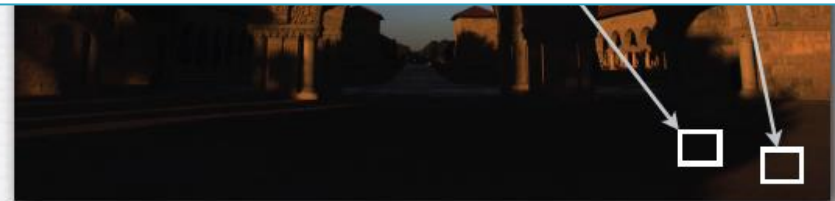


mean = 7

mean = 27



Based in these hints, calculate the solid angle under which we observe the Sun. (We assume the Sun is at the zenith.)

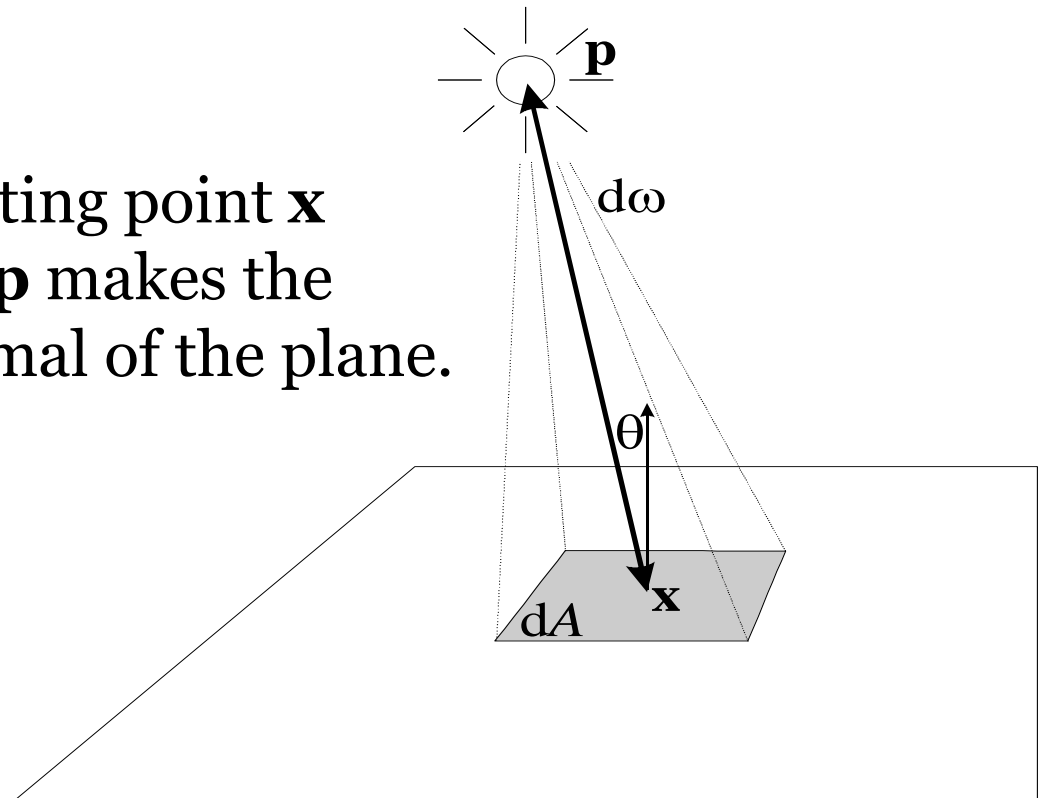


JPEG file

RAW, linearly boosted © 2009 Marc Levoy

Irradiance due to a point source

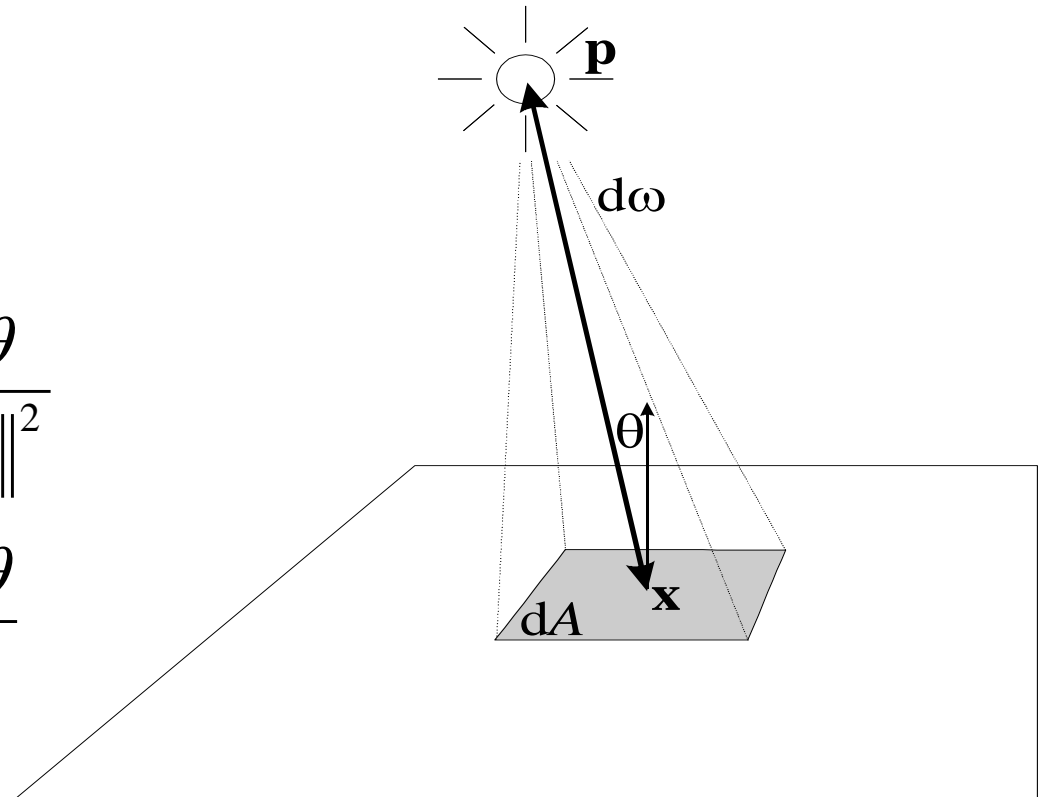
- What is the irradiance at point \mathbf{x} on a plane due to a point source with intensity $I(\omega)$ placed at the height h above the plane.
- The segment connecting point \mathbf{x} to the light position \mathbf{p} makes the angle θ with the normal of the plane.



Irradiance due to a point source

- Irradiance of a point on a plane lit by a point source:

$$\begin{aligned} E(\mathbf{x}) &= \frac{d\Phi(\mathbf{x})}{dA} \\ &= \frac{I(\mathbf{p} \rightarrow \mathbf{x})d\omega}{dA} \\ &= I(\mathbf{p} \rightarrow \mathbf{x}) \frac{\cos \theta}{\|\mathbf{p} - \mathbf{x}\|^2} \\ &= I(\mathbf{p} \rightarrow \mathbf{x}) \frac{\cos^3 \theta}{h^2} \end{aligned}$$



Area light sources

- Emission of an area light source is fully described by the emitted radiance $L_e(\mathbf{x}, \omega)$ for all positions on the source \mathbf{x} and all directions ω .
- The total emitted power (flux) is given by an integral of $L_e(\mathbf{x}, \omega)$ over the surface of the light source and all directions.

$$\Phi = \int_A \int_{H(\mathbf{x})} L_e(\mathbf{x}, \omega) \cos \theta \, d\omega \, dA$$

Diffuse (Lambertian) light source

- What is the relationship between the emitted radiant exitance (radiosity) $B_e(\mathbf{x})$ and emitted radiance $L_e(\mathbf{x}, \omega)$ for a Lambertian area light source?

**Lambertian source =
emitted radiance does not depend on the direction ω**

$$L_e(\mathbf{x}, \omega) = L_e(\mathbf{x}).$$

Diffuse (Lambertian) light source

- $L_e(\mathbf{x}, \omega)$ is constant in ω
- Radiosity: $B_e(\mathbf{x}) = \pi L_e(\mathbf{x})$

$$\begin{aligned} B_e(\mathbf{x}) &= \int_{H(\mathbf{x})} L_e(\mathbf{x}, \omega) \cos \theta \, d\omega \\ &= L_e(\mathbf{x}) \int_{H(\mathbf{x})} \cos \theta \, d\omega \\ &= \pi L_e(\mathbf{x}) \end{aligned}$$

Uniform Lambertian light source

- What is the total emitted power (flux) Φ of a **uniform** Lambertian area light source which emits radiance L_e
 - Uniform source – radiance does not depend on the position, $L_e(\mathbf{x}, \omega) = L_e = \text{const.}$

Uniform Lambertian light source

- $L_e(\mathbf{x}, \omega)$ is constant in \mathbf{x} and ω

$$\Phi_e = \mathbf{A} \mathbf{B}_e = \pi \mathbf{A} L_e$$